

Wavefront Sensor Comparison Testbed

Wavefront Sensing in the VLT/ELT Era V - AO Workshop Week II

13 - 15 October 2020

Mala Mateen Starfire Optical Range AFRL/RDSS



Collaborators

AFRL

Terry Brennan (Prime Plexus LLC) Robert Johnson Colton Bigler Scott Newey (Boeing) Jack Drummond

University of Arizona

Lauren Schatz Jared Males

Hart Scientific International

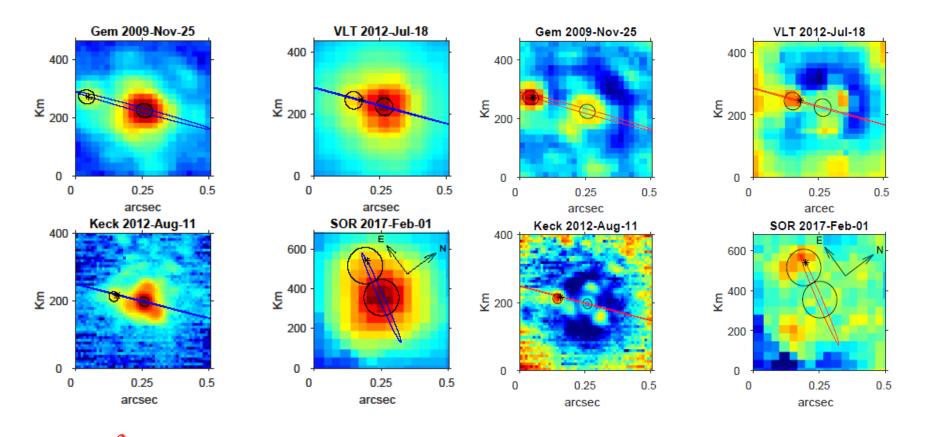
Michael Hart Johanan Codona



Overview

- Comparing the performance of the SHWFS, nICWFS, PWFS4, PWFS3
- Theory
- Simulations
- Laboratory demonstration

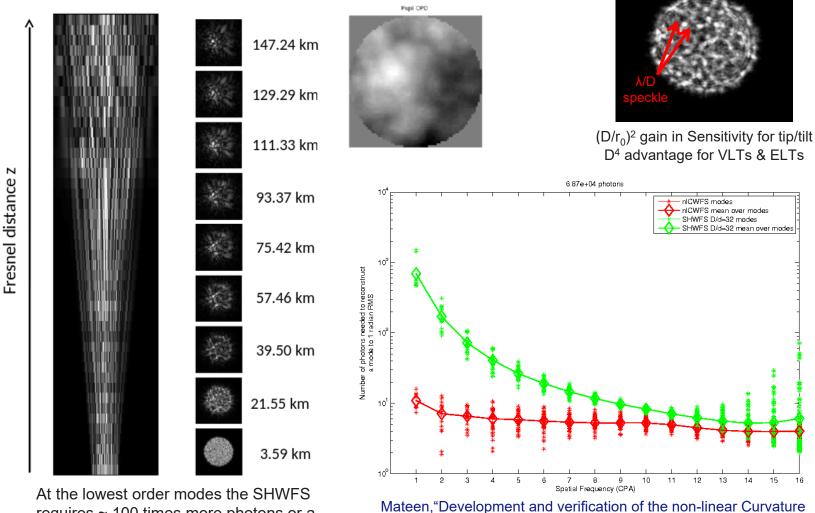
Objective: moons around asteroids & other closely spaced objects



Roxanne V = 13.0 and its moon at ΔJ = 3.0; Notice the competing aberrations - LWE. We want to obtain $\Delta J \ge 6.0$ at similar separations.

SUBMITTED Drummond & Colleagues: The orbit of asteroid (317) Roxane's satellite Olympias from Gemini, Keck, VLT and the SOR

Nonlinear Curvature Wavefront Sensor

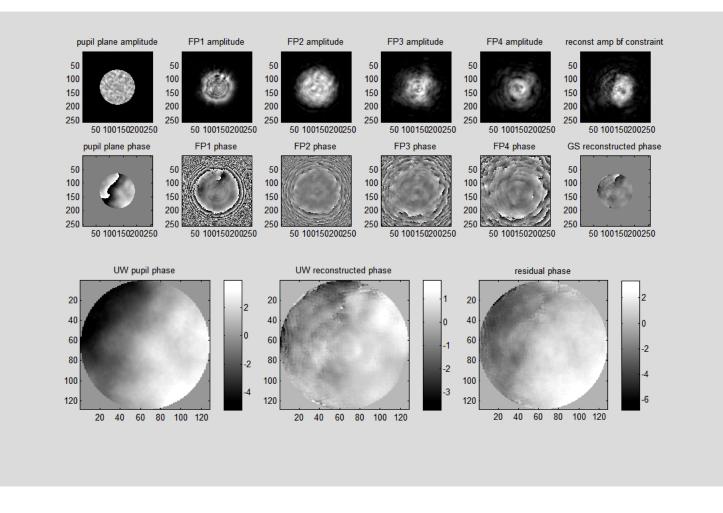


Wavefront Sensor" Ph.D thesis, University of Arizona (2015)

At the lowest order modes the SHWFS requires ~ 100 times more photons or a $\Delta m_v = 5$ to sense the mode.

THE AIR FORCE RESEARCH LABORATORY

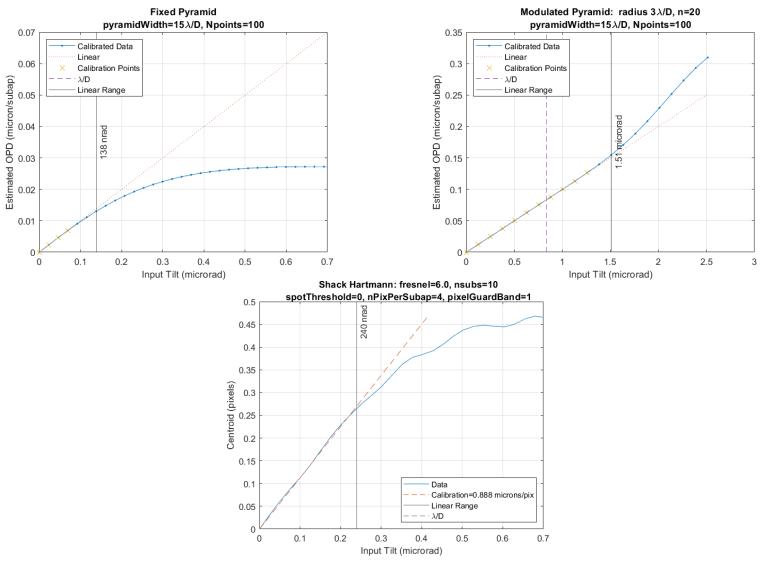
Gerchberg-Saxton based nICWFS Reconstruction



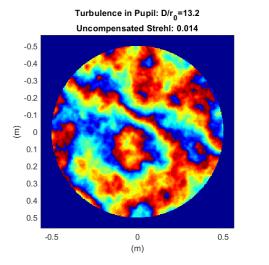
Faster reconstruction algorithm developed:

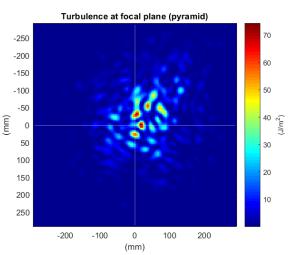
Codona et al. "A fast wavefront reconstructor for the nonlinear curvature wavefront sensor", Proc. SPIE 10703 (2018).

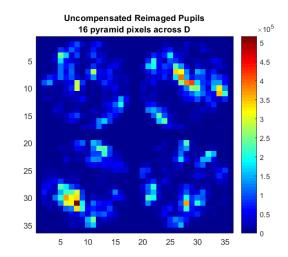
WFS Calibration Curves



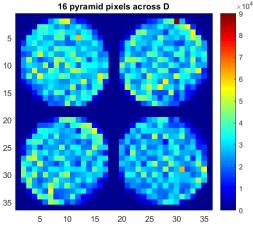
Pyramid Wavefront Sensor Simulation



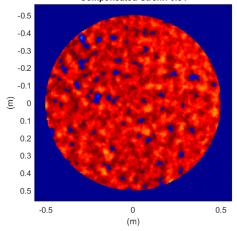




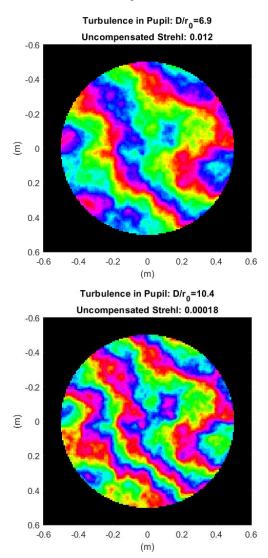
Compensated Reimaged Pupils 16 pyramid pixels across D

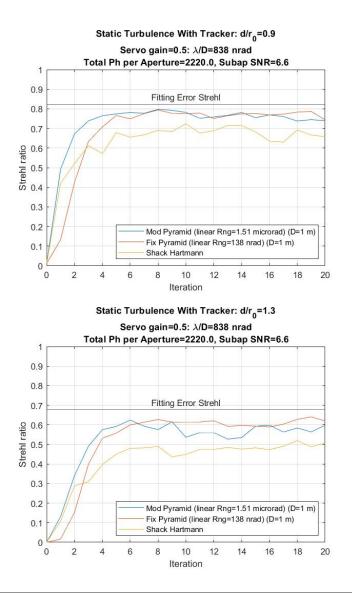


Modulated Pyramid Residual Phase Compensated Strehl: 0.84



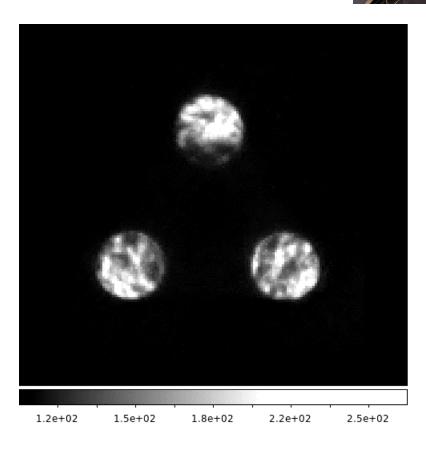
Strehl Comparison

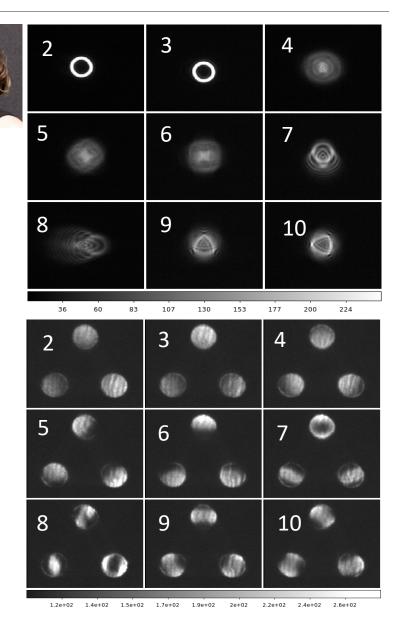






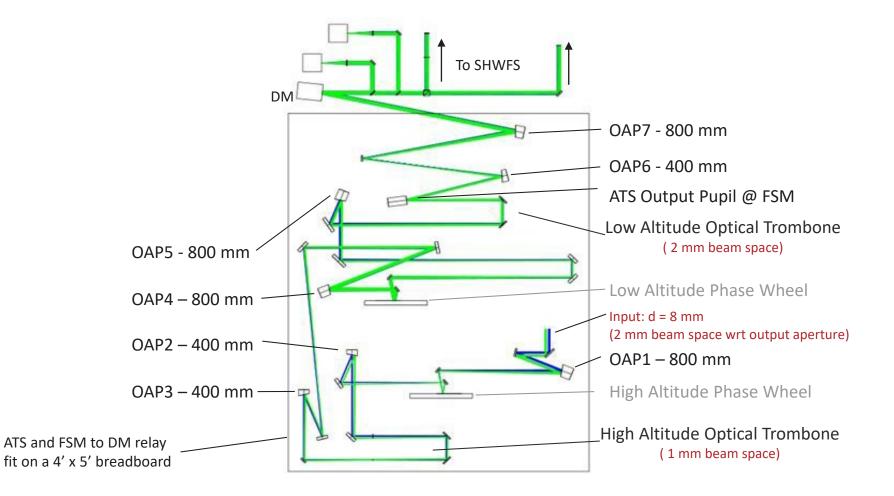
PWFS3 First Lab Results





Courtesy University of Arizona & Hart Scientific Inc.

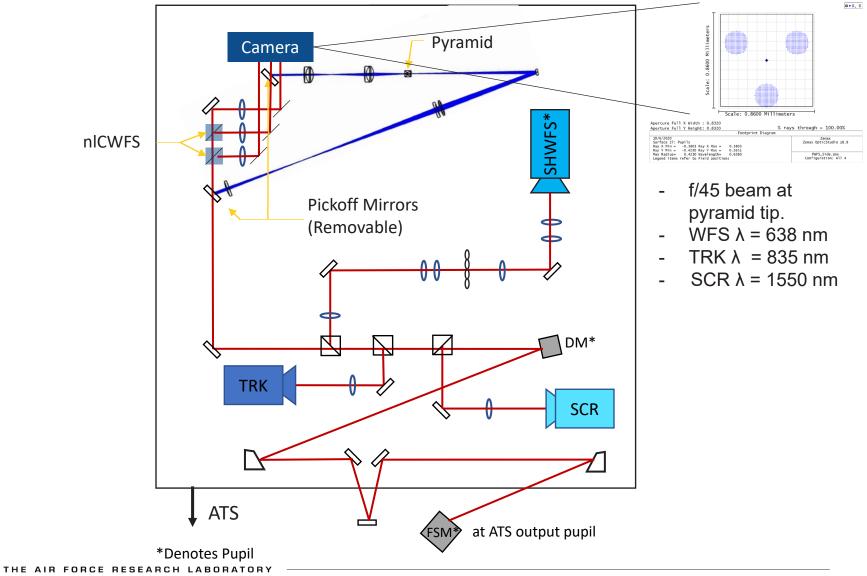
Atmospheric Turbulence Simulator Design



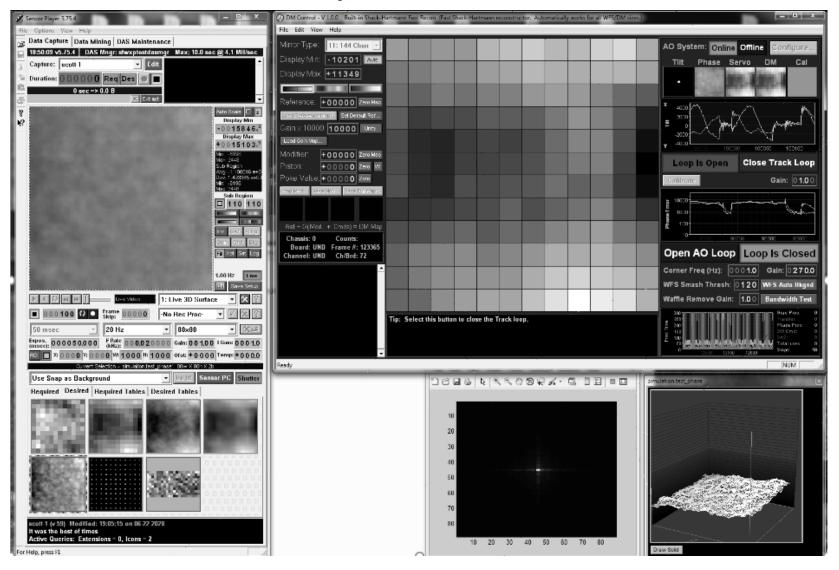
Phase wheel design is Guidestar Optical Systems Proprietary



WFS Comparison Testbed at Starfire Optical Range



Simulated AO Control System



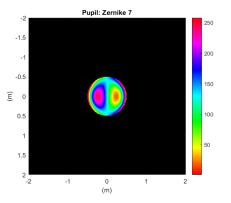


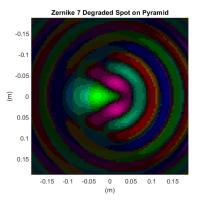
Questions?

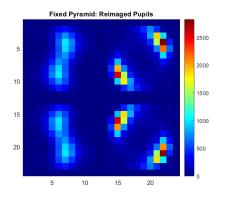


BackupanSlides

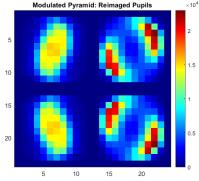
Zernike Test

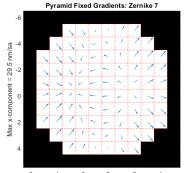




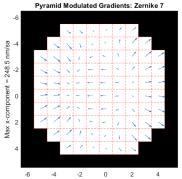


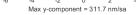
Modulated Pyramid: Reimaged Pupils

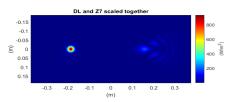


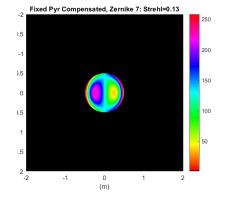


-6 -4 -2 0 2 4 Max y-component = 28.4 nm/sa

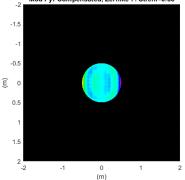




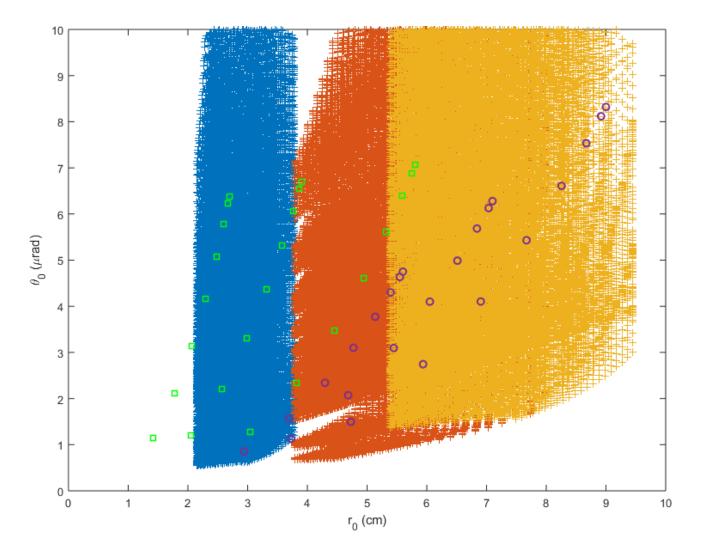




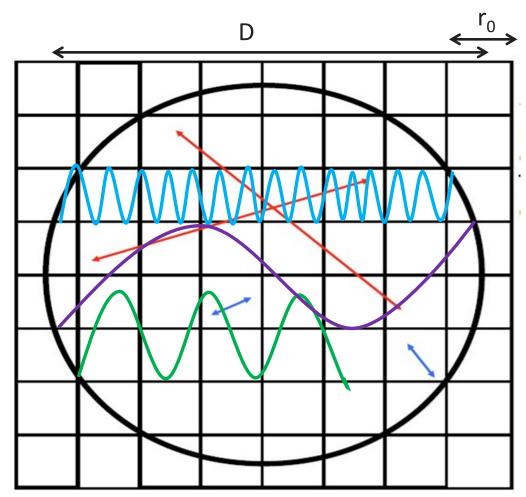
Mod Pyr Compensated, Zernike 7: Strehl=0.88



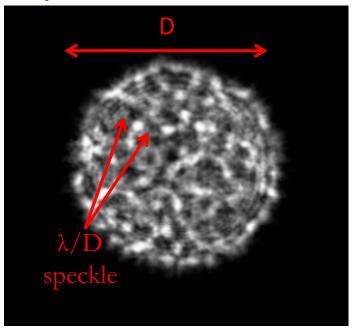
Atmospheric Turbulence Simulator Range



Extracting information from speckles



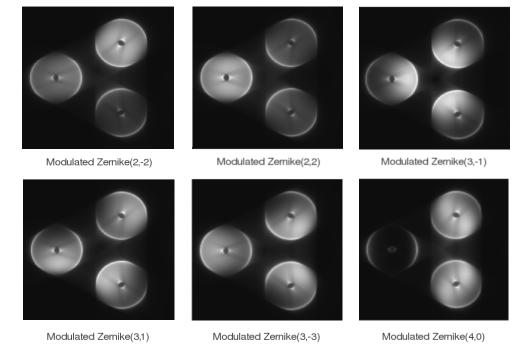
 $(D/r_0)^2$ gain in Sensitivity for tip/tilt



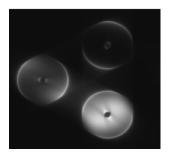
D⁴ advantage for VLTs & ELTs

SH subapertures: Seeing limited (λ/r_0)

PWFS3 Simulations

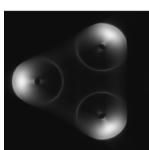


Modulated Zernike(1,-1)



Modulated Zernike(1,1)





Modulated Zernike(2,0)

Courtesy of Hart Scientific & University of Arizona